

# Non-Quantized Energy Values

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- ⊗ For the system of spin half particles the energy states were quantized
- ⊙ We will now generalize the discussion when energy is not necessarily quantized.
- ⊗ We will start with an example.

# N molecules in a box

→ Consider  $N$  gas molecules in a volume  $V$ .

→ The energy of the gas molecules will depend upon

② The potential energy of the  $N$  gas molecules determined the position  $r_1, r_2, \dots, r_N$  of the molecules

# Gas molecules in a Box

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- ①\* The kinetic energy of the molecules determined by their respective momenta;  $p_1, p_2, \dots, p_N$
- ②\* The internal energy of each molecule characterized by the internal configuration coordinates  $Q_1, \dots, Q_M$  and momentum coordinates  $P_1, \dots, P_M$ .

# Density of states

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→ Suppose  $\rho(E)$  denotes the density of states, which is the number of states that the system can take when the total energy is between  $E$  and  $E + \delta E$ .

# Density of states

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→ clearly  $E + \delta E$

$$\Omega(E) \propto \int \int \int_E dx_1 dx_2 \dots dx_N dp_1 \dots dp_N dq_1 \dots dq_N$$

where we integrate over all of the space that satisfy the constraint of total energy between  $E$  and  $E + \delta E$

# Ideal gas approximation

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→ Suppose the gas molecules do not have any "interaction energy".

Then

$$\begin{aligned} \Omega(E) &\propto \int_{E+\delta E}^E \int \int \dots dp_1, \dots, dp_N, da_1, \dots, da_N, dp_1, \dots, dp_N \\ &\quad \int_{r_i} \dots dr_1, \dots, dr_N \\ &= \int_{E+\delta E}^E \int \int \dots dp_1, \dots, dp_N, da_1, \dots, da_N, dp_1, \dots, dp_N \\ &\quad \int dr_1 \int dr_2 \dots \int dr_N \end{aligned}$$

# Towards Ideal gas

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$$\begin{aligned} & \frac{E + \delta E}{E} \\ & = \int \int_{E} dp_1 \dots dp_N d\theta_1 \dots d\theta_N dp_1 \dots dp_N \\ & \quad (V^N) \end{aligned}$$

→ Suppose, the molecules do not have any internal degrees of freedom and their internal energy  $E_{int} = 0$ .

Then

$$\Omega(E) \propto V^N \left[ \int \int_{E} dp_1 \dots dp_N \right]$$

# Limit of an ideal gas.

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$$\rightarrow \Omega(E) \propto V^N \int_{E}^{E+\delta E} \prod_{i=1}^N dp_{i1} \dots dp_{iN}$$

Suppose each molecule  $i$  with momentum  $\vec{p}_i = p_{i1} \vec{e}_x + p_{i2} \vec{e}_y + p_{i3} \vec{e}_z$

has energy  $\epsilon_i = \sum_{j=1}^3 \frac{1}{2m} p_{ij}^2$

$\rightarrow$  Total Energy is  $E = \sum_{i=1}^N \sum_{j=1}^3 \frac{1}{2m} p_{ij}^2$ .



# Limit of an ideal gas

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$$\rightarrow \therefore \Omega(E) \propto V^N \int_{\bar{E}}^{E+\delta E} \dots \int dp_1 \dots dp_N$$

and  $\int_{\bar{E}}^{E+\delta E} \dots \int dp_1 \dots dp_N$  is the

measure of the set

$$\left\{ (p_1, p_2, \dots, p_N) \mid E < \sum_{i=1}^N \sum_{j=1}^3 \frac{p_{ij}^2}{2m} \leq E + \delta E \right\}$$

which is the volume of a shell in  $3N$  dimensional space.

# Limit of an Ideal Gas

$$\Omega(E) \propto V^N \int \int \dots \int_{E}^{E+\delta E} dp_1 \dots dp_N$$

$\chi(E)$

$$\rightarrow \chi(E) = \int \int \int_{E}^{E+\delta E} dp_1 \dots dp_N$$

$$= \mu \left\{ p_{i_j} : i=1 \dots N, j=1 \dots 3; E \leq \sum_i \sum_j \frac{p_{i_j}^2}{2m} \leq E+\delta E \right\}$$

$$= \mu \left\{ p_{i_j} : i=1 \dots N, j=1 \dots 3; (2mE) \leq \sum_{i,j} \frac{p_{i_j}^2}{2m} \leq 2m(E+\delta E) \right\}$$



# Limit of an ideal gas

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$$\Rightarrow \Omega(E) \propto V^N \chi(E)$$

$\chi(E)$  is the volume of a shell of an hypersphere between the radius

$$R(E) \doteq \sqrt{2mE} \quad \text{and} \quad R(E+\delta E).$$

→ Volume of a hypersphere <sup>of dimension  $f$</sup>  with radius  $R(E) \propto (R(E))^f$

$$\Rightarrow \chi(E) = \left[ \frac{d}{dE} (R(E))^f \right] \delta E$$

# Limit of an ideal gas

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$$\Rightarrow \chi(E) = \int \frac{d}{dE} (2mE)^{f/2} \delta E$$

$$= \left( \frac{f}{2} (2mE)^{\frac{f}{2}-1} \right) \delta E$$

$$\Rightarrow \chi(E) \propto V^N \frac{3N}{2} (2mE)^{\frac{3N}{2}-1} \delta E; f=3N$$

$$\Rightarrow \chi(E) = B V^N E^{3N/2}$$

$B$  is a constant independent of  $N, V$ .

# Limit of an ideal gas

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→ Evidently

$\mathcal{U}(E)$  increases rapidly  
with  $E$

⊗ consider  $N \approx$  Avogadro's  
number

$$\mathcal{U}(E) = B V^N E^{3N/2}$$

# Generalization

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- Suppose the system has  $f$  degrees of freedom
- Let  $\phi_i(\epsilon)$  be the number of ways in which the  $i^{\text{th}}$  d.o.f. can have energy  $\leq \epsilon$
- $\Phi(E)$  be the number of ways in which the entire system can have energy  $\leq E$

# Density of states.

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Then  $\Phi(E) \leq \Phi_1(E/f) \Phi_2(E/f) \dots \Phi_f(E/f)$   
where the RHS is the number of ways in which the total energy is less than or equal to  $E$  with each d.o.f. having to satisfy the added constraint to have energy less than  $E/f$ .

→ Assume that  $\Phi_i(\epsilon) = \Phi_j(\epsilon)$



# Density of states

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$$- \Phi(E) \leq [\Phi_1(E/f)]^f$$

= We will assume

$$\Phi(E) \approx [\Phi_1(E/f)]^f$$

$\Omega(E)$  = Number of states of system between  
E and  $E + \delta E$

$$\begin{aligned} &= \frac{d\Phi(E)}{dE} \delta E = \delta E f \left[ \frac{\partial \Phi_1(E)}{\partial \varepsilon} \Big|_{\varepsilon=E/f} \right] \frac{1}{f} [\Phi_1(E/f)]^{f-1} \\ &= \left( \frac{\partial \Phi_1(E)}{\partial \varepsilon} \right)_{\varepsilon=E/f} [\Phi_1(E/f)]^{f-1} \delta E \end{aligned}$$

# Density of states

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$$\begin{aligned}\Rightarrow \Omega(E) &= \left. \frac{\partial \Phi_1(\varepsilon)}{\partial \varepsilon} \right|_{\varepsilon = E/f} [\Phi_1(E/f)]^{f-1} \delta E \\ &= \left( \left. \frac{\partial \Phi_1(\varepsilon)}{\partial \varepsilon} \right|_{\varepsilon = E/f} \delta E \right) [\Phi_1(E/f)]^{f-1}\end{aligned}$$

→ It is evident that a small increase in  $\Phi_1$  will lead to a large increase in  $\Omega$  due to the large exponent  $(f-1)$  in  $[\Phi_1(E/f)]^{f-1}$ . (Note that  $\partial \Phi_1 / \partial \varepsilon > 0$ ).

# Density of states

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$$\Rightarrow \ln \Omega(E) = (f-1) \ln \Phi_1 + \ln \left[ \frac{\partial \Phi_1}{\partial E} \delta E \right]$$

$$\approx f \ln \Phi_1 \quad \text{for } f \text{ large.}$$

If the energy  $E$  of the system is not too close to the ground state

[in which case  $\Phi_1(E) \approx 1$  and

$$\ln \Phi_1 \approx 0].$$

$$\Rightarrow \Omega(E) = \Phi_1^f \propto E^f$$

# Density of states

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→ Thus

$\Omega(E)$  is an EXTREMELY rapidly increasing function of energy.

# Thermal Interaction

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- In describing a macroscopic system it is in general possible to specify macroscopically measurable independent parameters  $x_1, x_2, \dots, x_n$  which are known to affect the equations of motion. (i.e. appear in the Hamiltonian)
- Applied magnetic field, volume of a system
- If a particular quantum state has energy  $E_r$ . then  $E_r = E_r(x_1, x_2, \dots, x_n)$ .

# Macroscopic State

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- The macroscopic state or the macrostate is defined by specifying external parameters and any other constraints imposed on the system.
- Corresponding to a macrostate, the system can be in any one of a very large number of microstates

# Thermal Interaction

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- A and A' are two macroscopic systems
- $A^{(0)} = A \cup A'$  is isolated (closed) system.
- Suppose all possible external parameters remain fixed; thus, the energy levels of the system A and system A' remain fixed. [External magnetic field of spin 1/2 example]

# Thermal Interaction

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→ Every system  $A$  and  $A'$  that are prepared in accordance with the specification will have energy transfer with different  $(A, A')$  elements having different energy transfer.

→ Mean Energies can be discussed.

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# Heat

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- The mean Energy transferred from one system to the other as a result of purely thermal interactions is called heat
- The change  $\overline{\Delta E}$  of the mean energy of system A is called the heat  $Q$  absorbed by the system.
- $$Q \equiv \overline{\Delta E}$$

# Conservation

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→ Since Energy of  $A^{(s)}$  is conserved

$$\overline{\Delta E} + \overline{\Delta E'} = 0$$

$\overline{\Delta E'}$  is the mean change in energy of  $A'$ .

$$\therefore Q + Q' = 0 \quad \text{or} \quad Q = -Q'$$

# Thermal Equilibrium.

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→ Suppose A had energy  $E_i$  initially and A' had energy  $E'_i$ . Then

$$E + E' = E'_i + E_i = E^{(0)} = \text{constant}$$

as  $A^{(0)}$  is closed.

→ In thermal equilibrium all <sup>accessible</sup> states of the closed system are equally likely.

# Most probable configuration

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→ Number of accessible states of A

If energy of A is between  $E$  and  $E + \delta E$   
 $= \Omega(E)$ .

→ For each accessible state of A with energy  $E$ , A' will have  $\Omega'(E')$  with  $E' = E^{(0)} - E$  accessible states

→ Total number of accessible states for  $A^{(0)} \propto \Omega(E) \Omega'(E')$

# Most probable configuration

→ As all <sup>accessible</sup> states are equally likely the most probable energy configuration is the one that has the most number of accessible states which is obtained by setting

$$\frac{\partial}{\partial E} [\Omega(E) \Omega'(E')] = \frac{\partial}{\partial E} \Omega(E) \Omega'(E-E') = 0$$

which has the same solution as

$$\frac{\partial}{\partial E} (\ln \Omega(E) \Omega'(E-E')) = \frac{\partial}{\partial E} \ln \Omega(E) + \frac{\partial}{\partial E} \ln \Omega'(E') = 0$$

# Most probable energy configuration

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$$\Rightarrow \frac{\partial}{\partial E} \ln \Omega(E) - \frac{\partial}{\partial E'} \ln \Omega'(E') = 0$$

$\Rightarrow$

$$\frac{\partial}{\partial E} \ln \Omega(E) = \frac{\partial}{\partial E'} \ln \Omega'(E')$$

$\rightarrow$  Defining  $\sigma(E) = \ln \Omega(E)$

we have  $\frac{\partial \sigma(E)}{\partial E} = \frac{\partial \sigma'(E')}{\partial E'}$

# Most probable Energy Configuration

→ Let  $\beta = \frac{\partial \ln \Omega}{\partial E}$  and  $\sigma = \ln \Omega$ .

$$-kT = 1/\beta \quad ; \quad \tau = 1/\beta.$$

$$S = k \ln \Omega$$

$$\Rightarrow \frac{1}{T} = k\beta = k \frac{\partial \ln \Omega}{\partial E} = \frac{\partial S}{\partial E}$$

Thus, the most probable energy configuration occurs when  $S+S'$  is maximized or when  $T = T^*$ .